



Presented at the COMSOL Conference 2009 Boston

Using Microwaves for Extracting Water from the Moon

COMSOL Conference
October 9, 2009
Boston, MA

Edwin Ethridge, PhD, Principal Investigator
Materials & Processes Laboratory
NASA-MSFC-EM40, Huntsville, AL

William Kaukler, PhD
Department of Chemistry
The University of Alabama in Huntsville

Frank Hepburn
Materials & Processes Laboratory
NASA-MSFC-EM20, Huntsville, AL



Water is one of the Most Plentiful Compounds in the Universe

- **Earth's water came from comets early in its history.**
- **Mars has vast quantities of water not only at the poles but also at lower latitudes ($>55^{\circ}$).**
- **Significant quantities of water are present on several moons of Jupiter (Europa, Ganymede, and Callisto), Saturn (Enceladus), and Neptune (Triton).**

Chronology of “Water on the Moon”



1959, prior to Apollo, scientists **speculated** that there should be some residual **water on the moon from cometary impacts**.

Apollo found “**no water**”.

1994, the SDI-NASA **Clementine spacecraft** mapped the surface. **Microwave radio signals from South pole shadowed craters were consistent with the presence of water**.

1998 – Prospector - Neutron Spectrometer – high H concentrations at poles

September 2009 – water observed at diverse locations of moon

Chandrayaan-1 - Moon Mineralogy Mapper (M3)

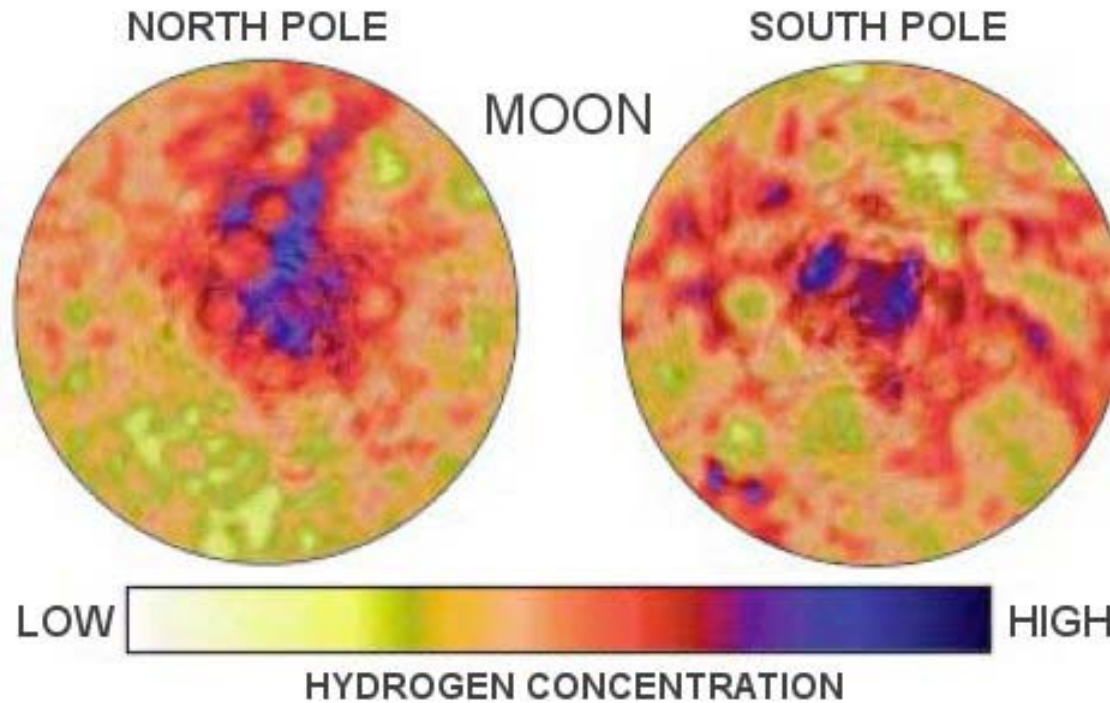
Cassini - Visual and Infrared Mapping Spectrometer (VIMS)

EPOXI spacecraft - High-Resolution Infrared Imaging Spectrometer

October 9, 2009 - Lunar CRater Observation and Sensing Satellite - LCROSS



1998- Lunar Prospector Hydrogen Maps of the Lunar Poles



1998 Lunar Prospector - Neutron Spectrometer scanned for hydrogen-rich minerals. **Polar craters** yielded **neutron ratios** which indicated **hydrogen => H₂O**

Average H₂O concentration ~2% → Theoretically approaching Billions of tons



Lunar CRater Observation and Sensing Satellite - LCROSS

- **THIS MORNING** the **Centaur upper stage impacted** a permanently shadowed crater, Cabeus A near the south pole of the Moon.
- **Mission Objective - confirm the presence of water ice** in a permanently shadowed crater at the **Moon's South Pole**.
- **Spectral analysis** of the resulting impact plume will look for **water ice**.

Crater Cabeus A →

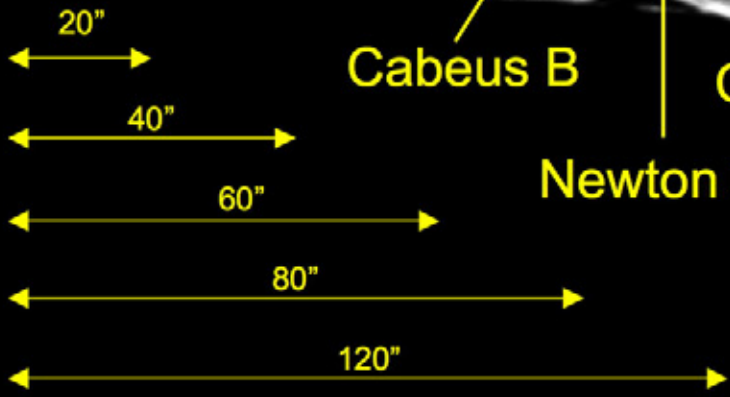
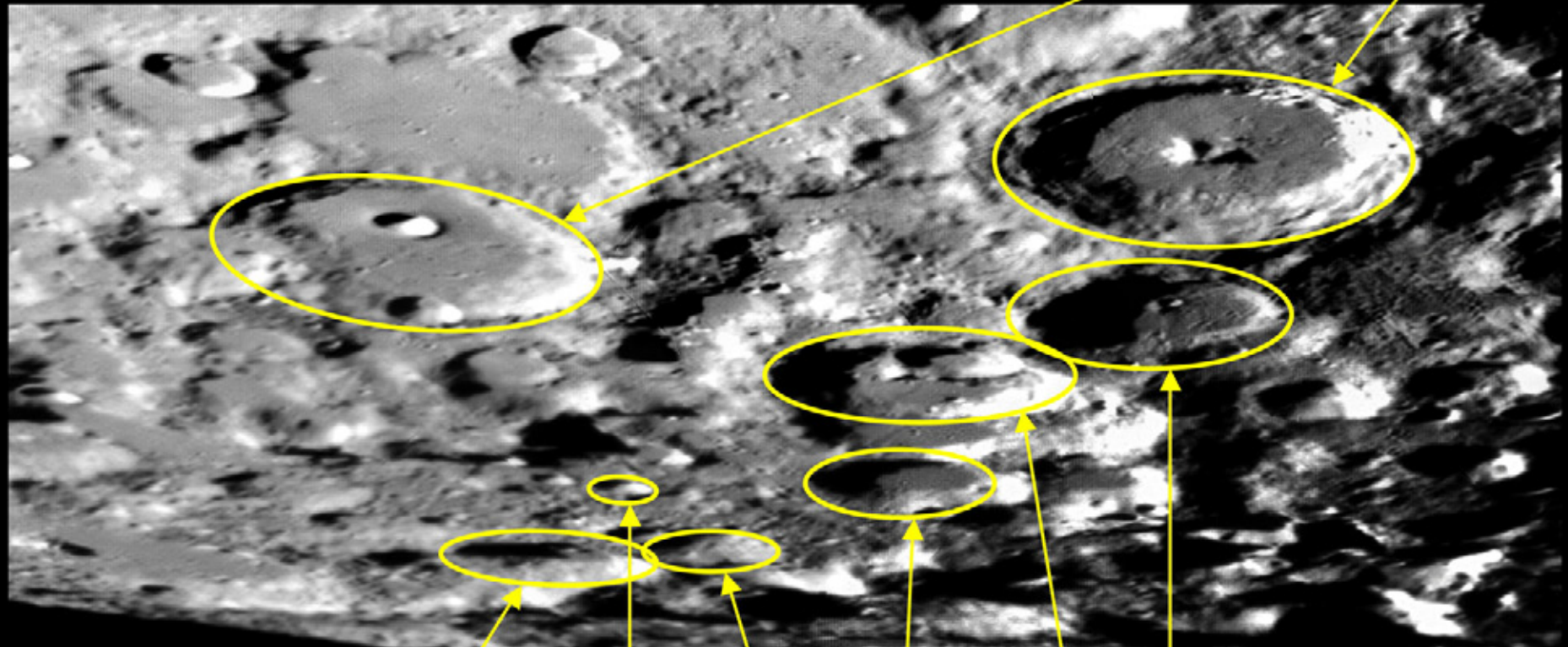


Lunar South Pole Map

Medium field

Casatus

Moretus



Cabeus B

Cabeus A

Newton E

Newton A

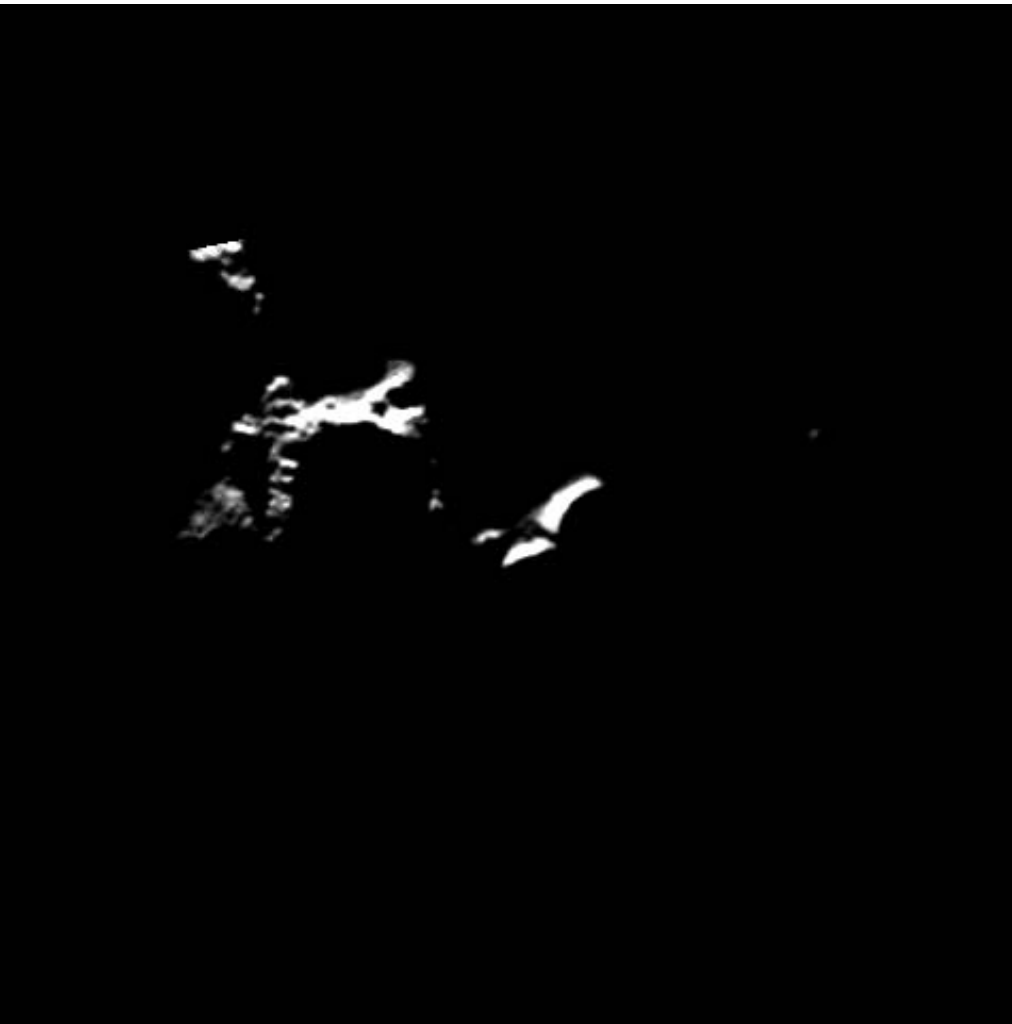
Newton

Short

NMSU / MSFC
Tortugas Observatory 24"
0.9 - 1.7 μm InGaAs Camera

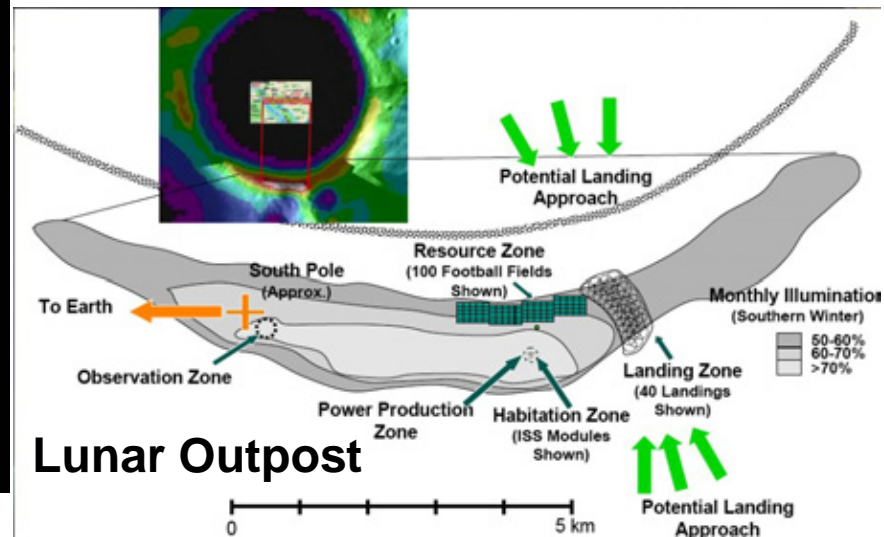
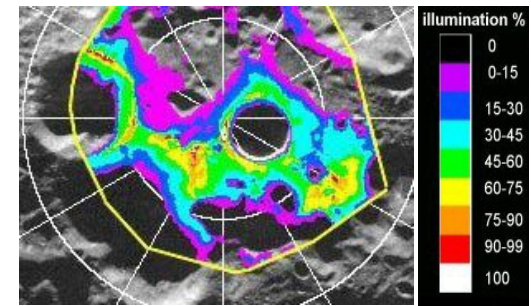
Scale at time of LCROSS impact (1.78 km/arcsec)

Illumination of the South Polar Region of the Moon over one Lunar Day (28 days)



<http://www.lpi.usra.edu/resources/clemen/clemen.html>

Moon Polar light map



Continuous Sunlight, Continuous Contact with Earth, “Moderate” Temperature, and Resources (water ice?)

Importance of Water

- **Water (and oxygen) for the manned lunar outpost.**
Resupply requirements:
1 ton water and 1 ton oxygen per year
~\$10,000 to \$100,000/ pound from Earth to moon
Increasing to 10 tons per year
- **Eventually (perhaps) - rocket propellant for exploration,**
Electrolyze water into H and O,
liquify to LOX and LH
10's of tons not launched out of Earth's
gravitational well.

Water on the Moon

The water on the moon is in a number of different forms.

Chemically Bound water – Hydroxyl groups -OH

- **The hydroxyls can be on the surface of the grains of soil.**
- **They can also be part of the chemical composition of the rock.**
- **This could include other hydrated chemical species (at the poles).**
- **Relatively small weight percent measured in Apollo soils.**

Molecular water H₂O

- **Recently measured water vapor near the lunar surface.**
- **Water molecules lightly bound to Si⁺ dipoles on the surface of grains of lunar soil.**
- **Water ice physically condensed at poles. This physically condensed cryogenically trapped water ice is speculated to be present in relatively high concentrations (on the order of 2 weight percent),**
This is the water we have proposed to extract with microwaves.

Use of Microwaves for the Extraction of Lunar Water



- **Lunar Soil (in vacuum) is a Super Thermal Insulator
Is like an aerogel, very low heat flow.**
- **Microwave energy penetrates the soil heating from the
inside out. Penetration depth is dependent
on Frequency and dielectric properties.**
- **Conversion from electricity through microwaves to heat,
efficient, heating causes sublimation of water ice.**
- **Excavation may not be required,
Cryogenic water ice is as hard as granite
Saving energy, infrastructure, and equipment
Little if any disruption of lunar dust (hazard)**

“Moon in a Bottle” Laboratory Proof of Principle

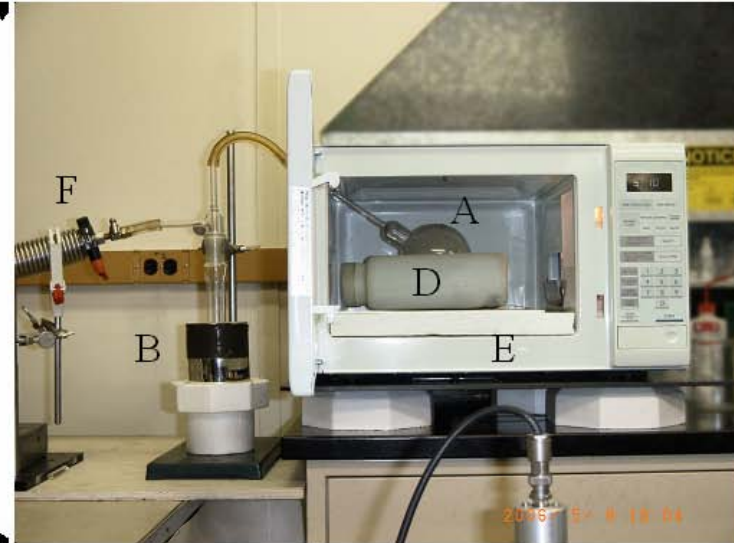
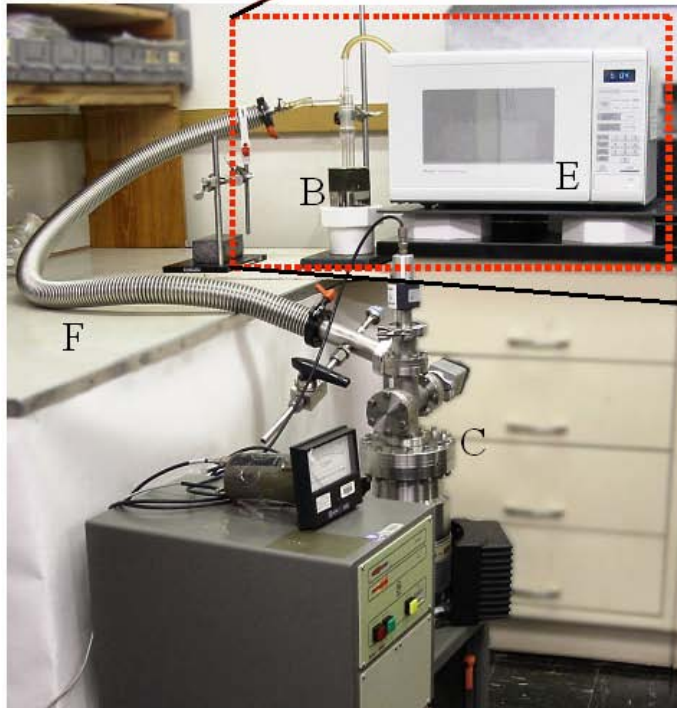


Fused silica vessel with lunar permafrost simulant.



Experimental Facilities

- Standardized lunar regolith simulant (JSC-1A), particle size distribution and chemistry of (Apollo 14).
- Water ice concentration (2 weight %, 2g)
- Temperature (-196 to -50C), LN2,
- Vacuum level (10⁻⁵ torr),



- Bench top microwave facility**
- A. Vacuum quartz lunar regolith simulant vessel
 - B. Liquid nitrogen cold-trap
 - C. Turbo-molecular vacuum pump
 - D. LN2-cooled regolith simulant
 - E. Microwave oven chamber

Extraction Efficiency

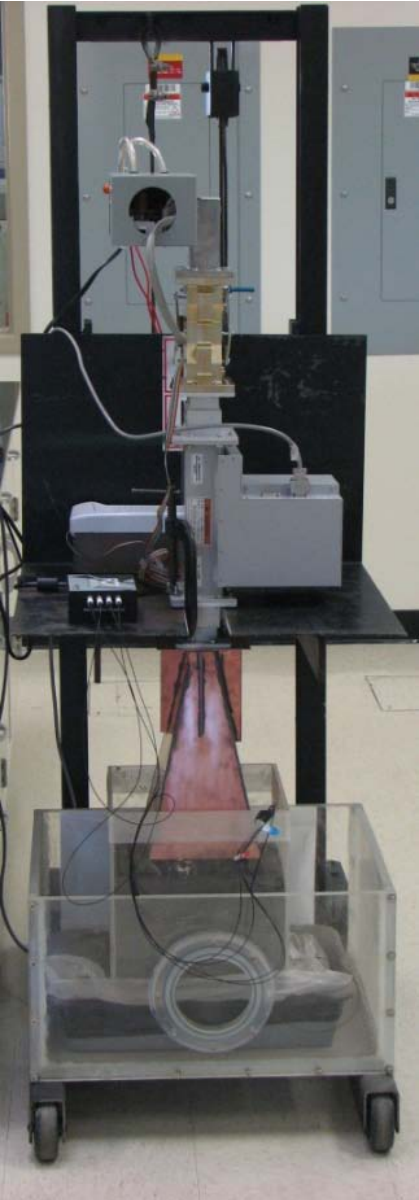


- **Microwaves coupled well to soil simulant at LN2 temperature.**
- **The regolith and the cold trap were weighed before and after the experiment.**
- **At least 95% of the water added to the regolith simulant was extracted (in 2 minutes) all below 0°C.**
- **Of the extracted water 99% was captured in the remote cold trap.**

Microwave Lunar Water Extraction Prototype

- Magnetron source (2.45 GHz, 1100 W) with isolator, auto-tuner and copper high-gain horn.
- Mounting provides mobility over surface and height adjustment of horn.
- Temperatures within the bed of simulant (JSC-1A) were made using fiber optic temperature sensor in place during heating.

- Vacuum chamber evaluation of the microwave penetration and water vapor --> permeability through lunar soil simulant.



Attenuation - Beer-Lambert law

Penetration Depth*

$$\text{Penetration} \sim \frac{\lambda (\epsilon')^{1/2}}{2 \pi \epsilon''}$$

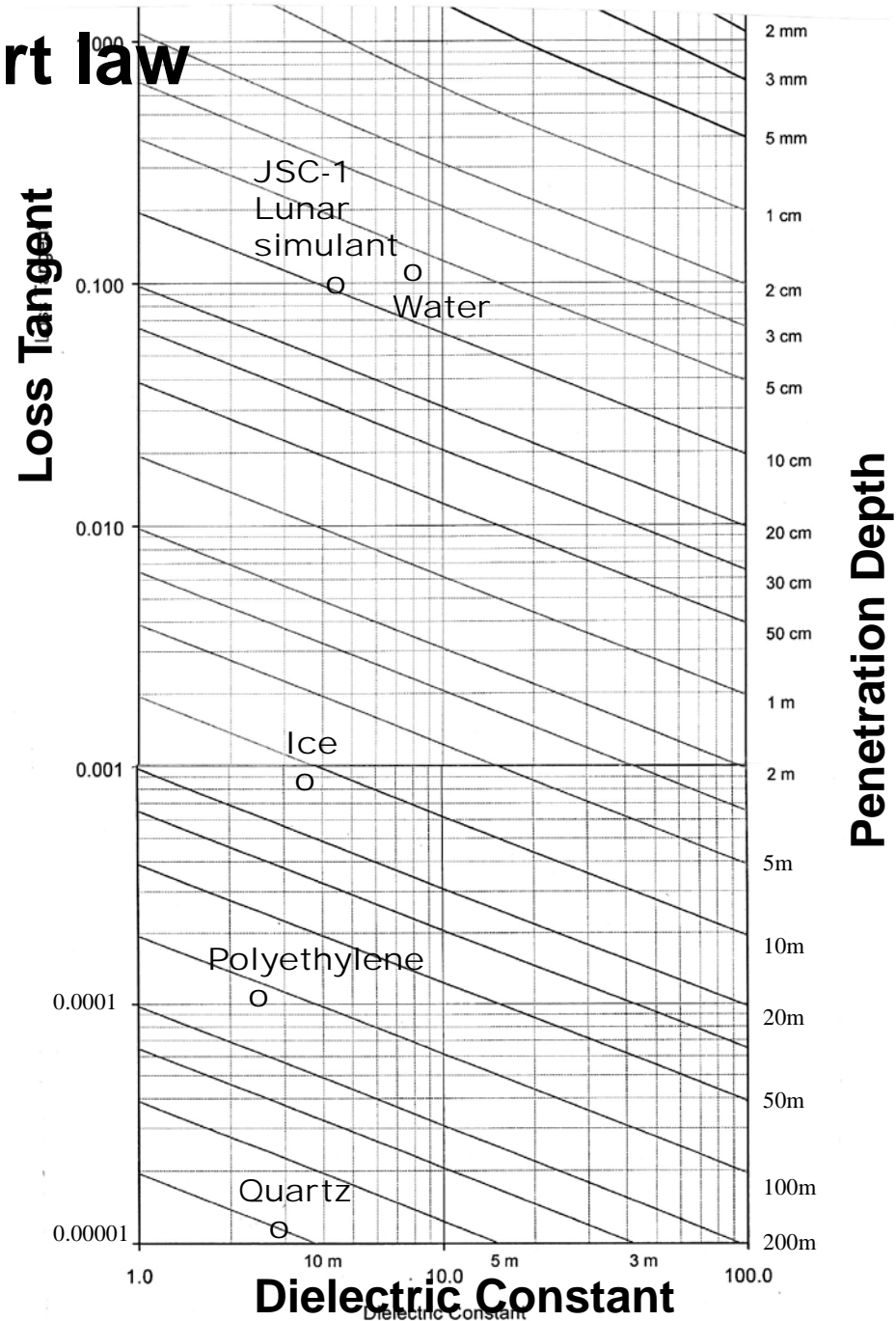
	ϵ''	ϵ'	Penetration
Water	0.1	8	6cm
Simulant JSC-1	0.1	4	10cm
Ice	0.001	3	1m
Polyethylene	0.0001	2	100m
quartz	0.00001	3	1000m

$$\epsilon = \epsilon' + j \epsilon''$$

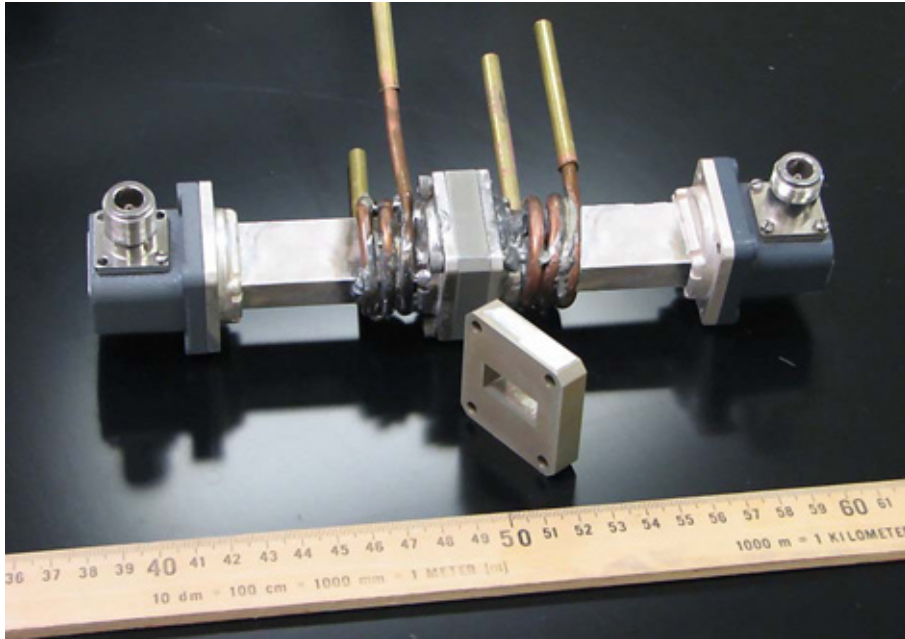
ϵ' dielectric constant

ϵ'' loss factor

*1/e (0.37) of the original value.



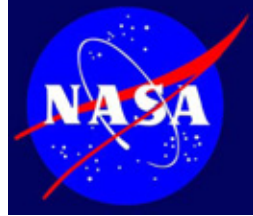
Dielectric Property Measurements Lunar Soil Simulant – JSC-1A



With Frank Hepburn - EM20

- Our custom fabricated 10 GHz (range 8 – 12 GHz) X-band waveguide apparatus for dielectric measurements over a range of temperatures, LN₂ to above room temperature.
- Heating coils near the coax connectors (not shown) keep the instrument connections at room temperature while the sample residing between the cooling the coil is chilled with free-flowing LN₂.

Dielectric Properties of JSC-1A Simulant X-band 10 GHz, room temperature



Real and imaginary

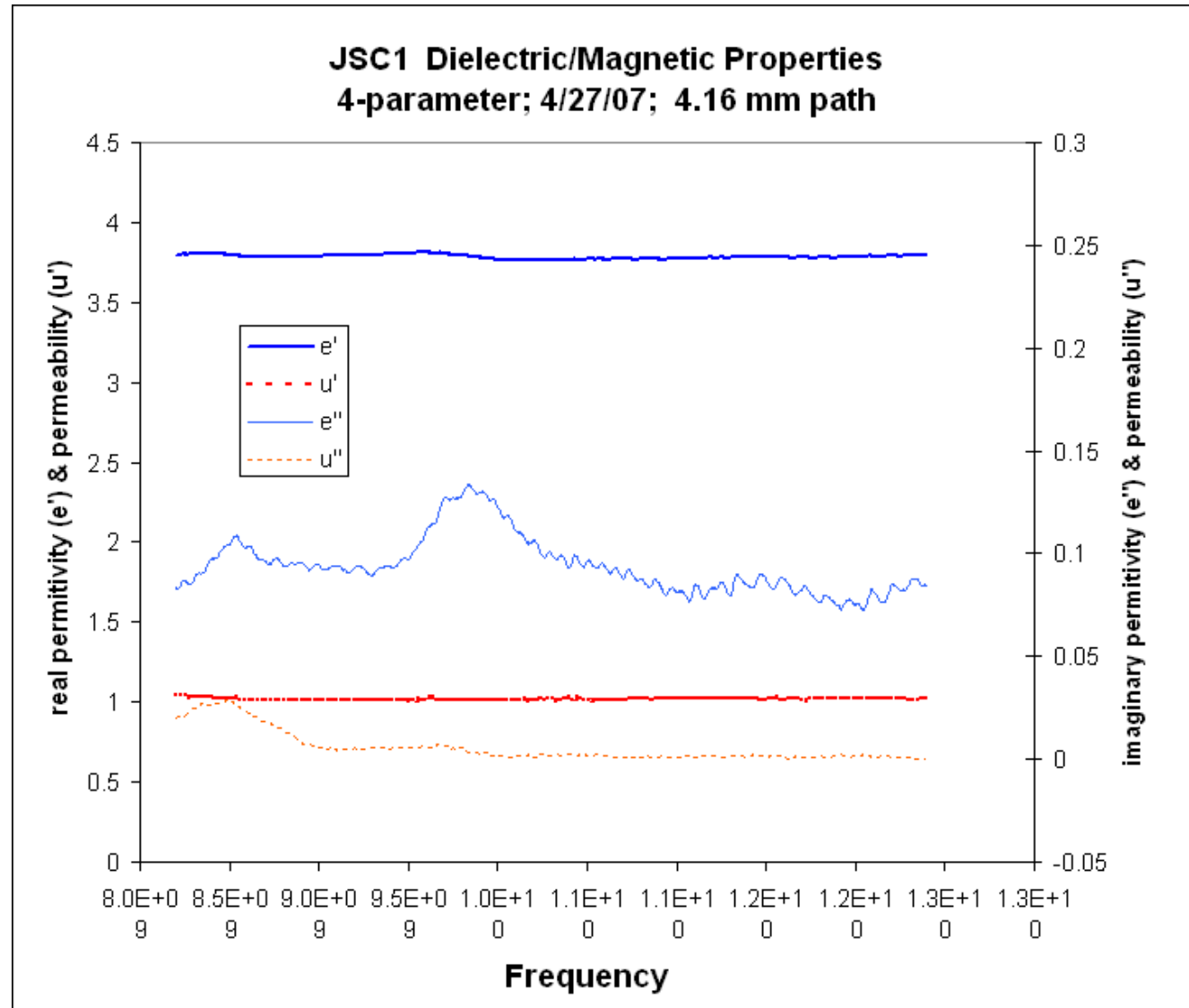
1. Electric Permittivity

(dielectric constant & Loss factor) and

2. Magnetic Permeability

We expect that Nano-phase Fe in lunar soil will significantly affect the permeability.

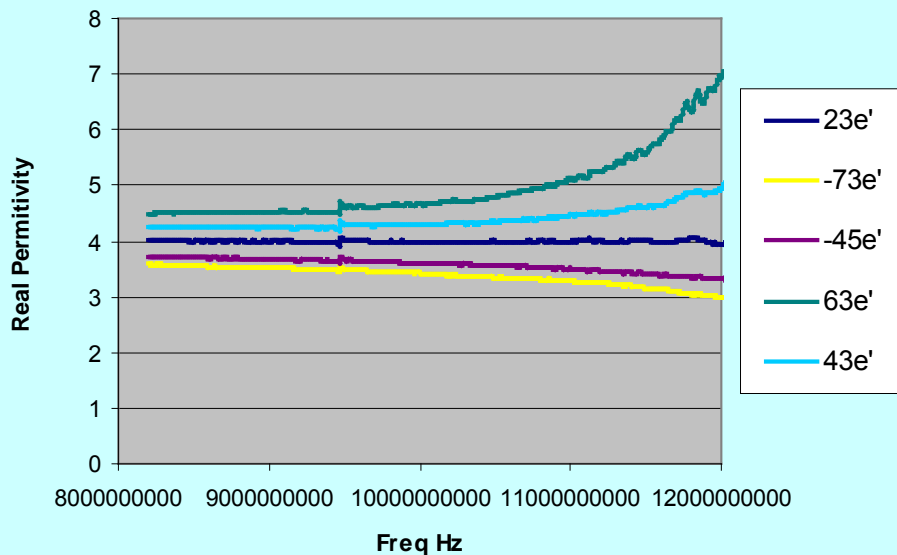
Proposal pending:
Loan of **Apollo Soil** sample to measure dielectric properties.



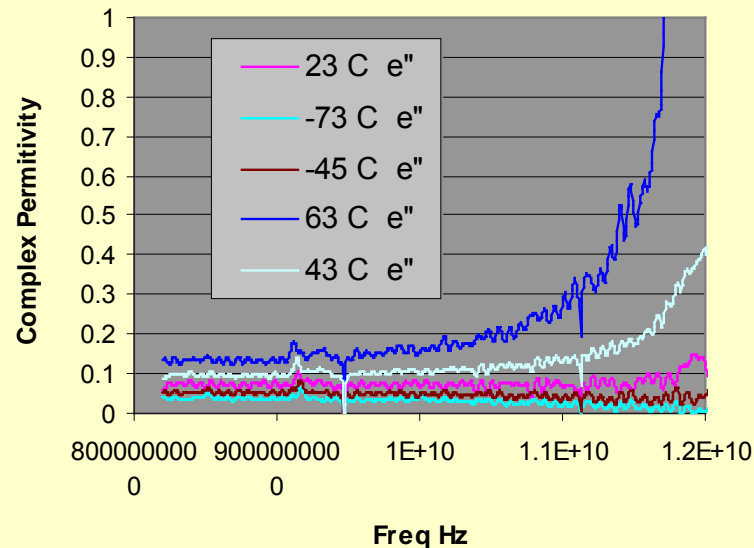
Temperature Dependence (-73C to 63C) of the Dielectric Properties of Lunar Soil Simulant JSC-1A



JSC-1A Permittivity vs Temperature
X-band, 8 to 12 GHz, Oct 30, 2007



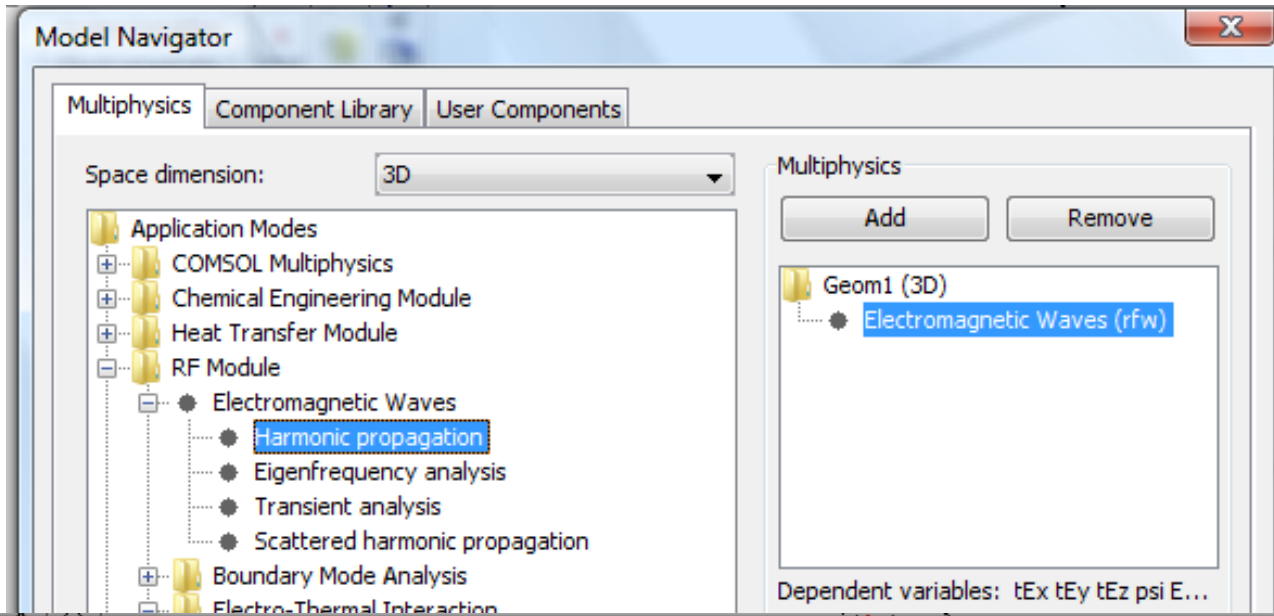
JSC-1A Complex Permittivity vs Temperature
X-band, 8 to 12 GHz, Oct 30, 2007



Dielectric constant (real component of permittivity) vs. frequency over the X-band (8 to 12 GHz) for JSC-1A, temperatures from 63 C to -73 C.

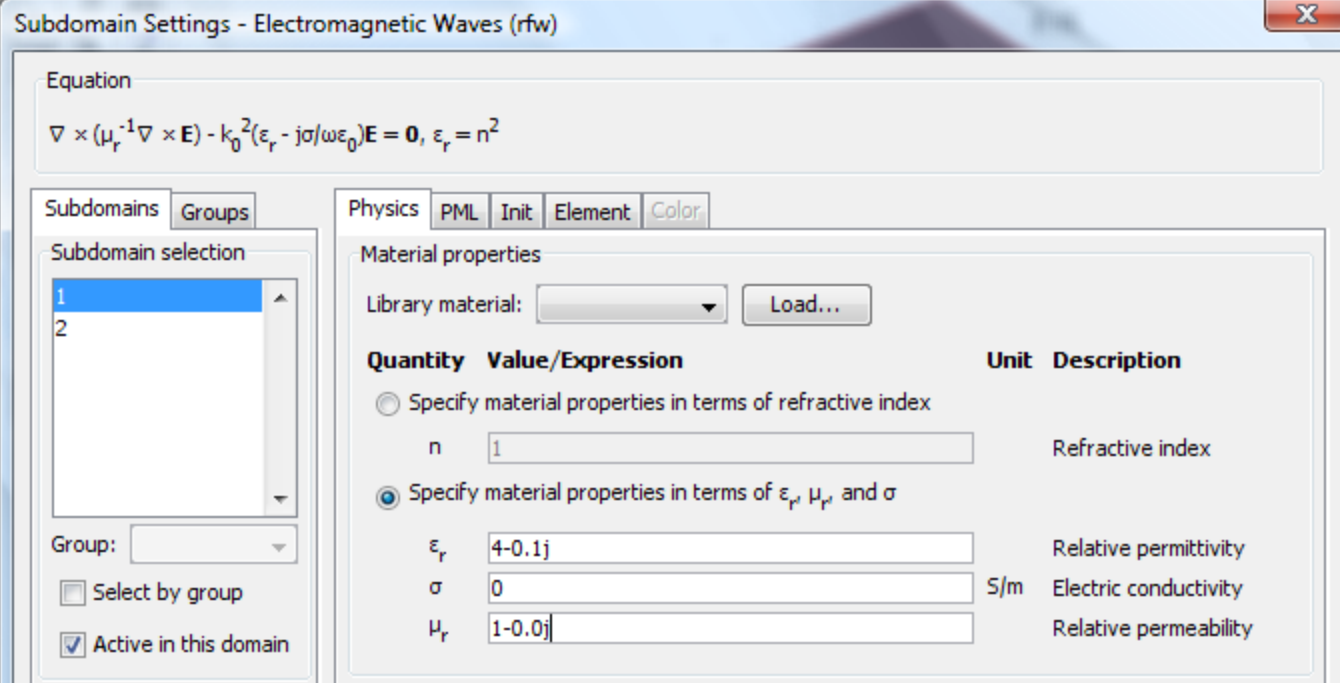
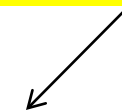
The loss factor (imaginary part of permittivity) vs. frequency 8 to 12 GHz 63 to -73 C.

COMSOL - RF Module – Electromagnetic Waves – Harmonic Propagation



COMSOL 3.5a

Subdomain Settings



Electric Permittivity

4 - 0.1j

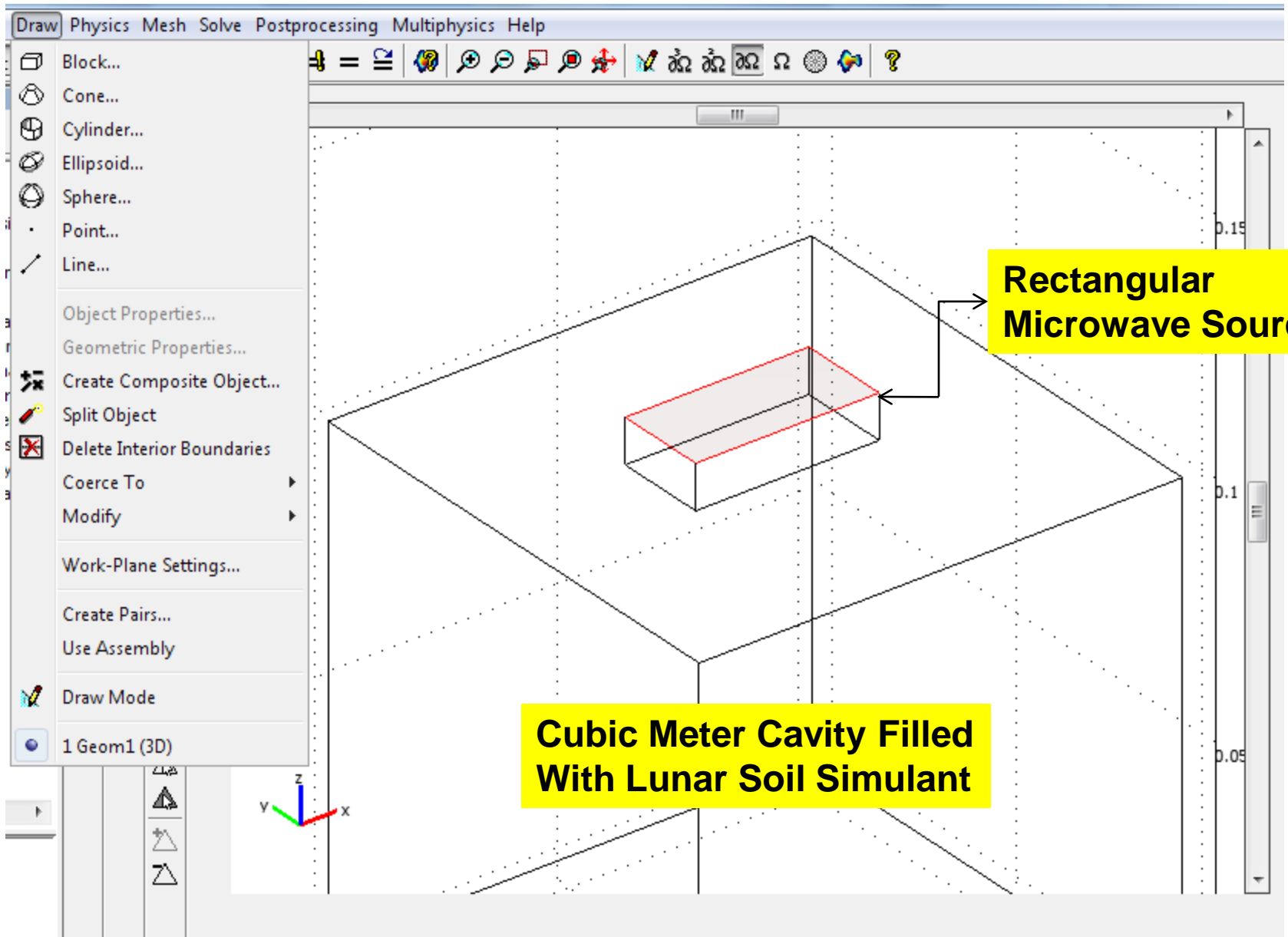
Magnetic Permeability

1 - 0.0j

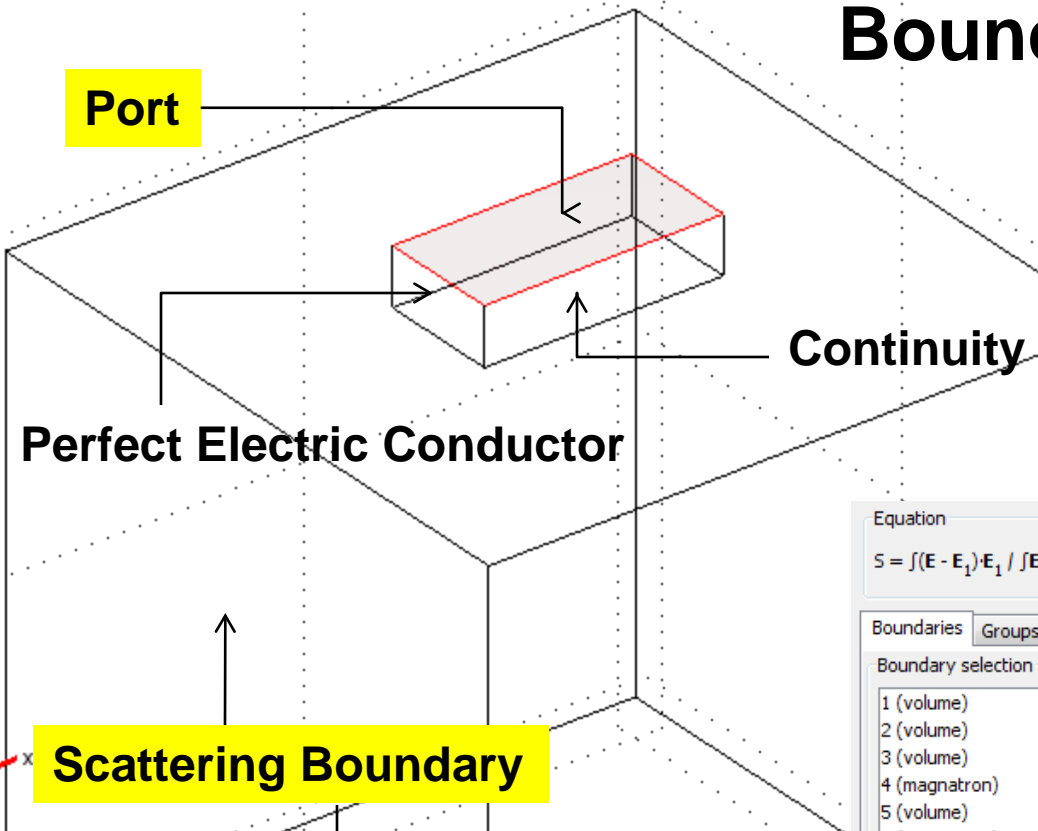
Lunar Soil Simulant Dielectric Properties



Geometry - Draw



Boundary Conditions



Rectangular Port

Conditions | Material Properties | Port | Far-Field | Color

Port definition

Mode specification: Rectangular

S-parameter output: Magnitude and phase

Mode type: Transverse electric (TE)

Mode number: 10

Port

Equation

$$S = \frac{\int (\mathbf{E} - \mathbf{E}_1) \cdot \mathbf{E}_1}{\int \mathbf{E}_1 \cdot \mathbf{E}_1}$$

Boundaries | Groups

Boundary selection

- 1 (volume)
- 2 (volume)
- 3 (volume)
- 4 (magnatron)
- 5 (volume)
- 6 (magnatron)
- 7 (magnatron)
- 8 (window)
- 9 (Port)

Conditions | Material Properties | Port | Far-Field | Color

Boundary sources and constraints

Boundary condition: Port

Port number: 1

Wave excitation at this port

Quantity	Value/Expression	Unit	Description
P_{in}	10	W	Port power level
Φ_p	0		Port phase

Equation

$$\mathbf{n} \times (\nabla \times \mathbf{E}) - j\mathbf{k} \times (\mathbf{E} \times \mathbf{n}) = -\mathbf{n} \times (\mathbf{E}_0 \times j\mathbf{k}(\mathbf{n} - \mathbf{k})) \exp(-j\mathbf{k} \cdot \mathbf{r})$$

Boundaries | Groups

Boundary selection

- 1 (volume)
- 2 (volume)
- 3 (volume)
- 4 (magnatron)
- 5 (volume)
- 6 (magnatron)
- 7 (magnatron)
- 8 (window)
- 9 (Port)
- 10 (magnatron)

Conditions | Material Properties | Port | Far-Field | Color

Boundary sources and constraints

Boundary condition: Scattering boundary condition

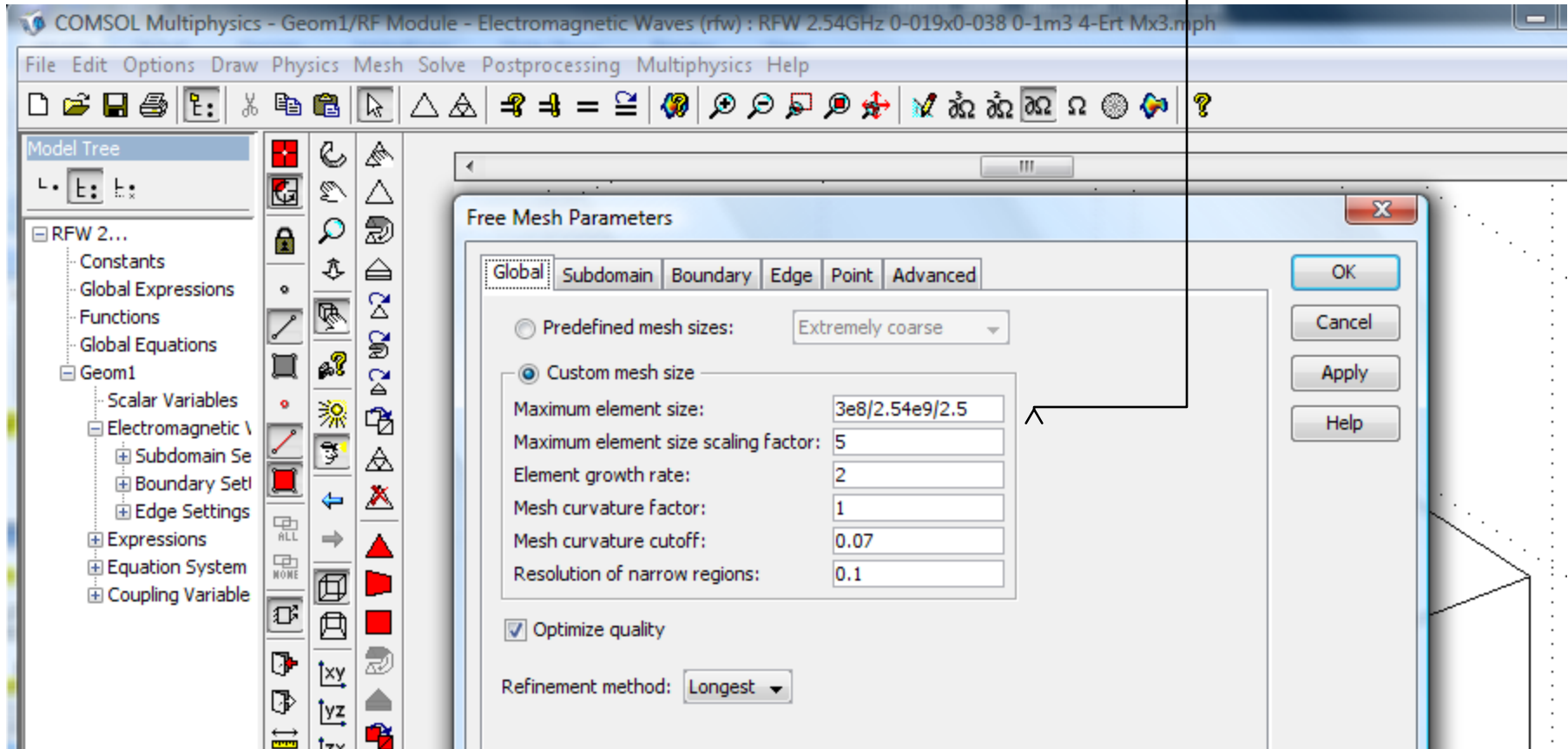
Quantity	Value/Expression	Unit	Description
Incident field:	Wave given by E field		
\mathbf{E}_0	0 0 0	V/m	Electric field
Wave type:	Plane wave		
\mathbf{k}	-nx_rfw -ny_rfw -nz_rfw		Wave direction

Port Power 10 Watt

Meshing must satisfy the Nyquist criteria

$$\begin{aligned}\text{Maximum element size} &= c / \lambda / 2.5 \\ &= 3e8 / 2.45e9 / 2.5\end{aligned}$$

Free Mesh Parameters

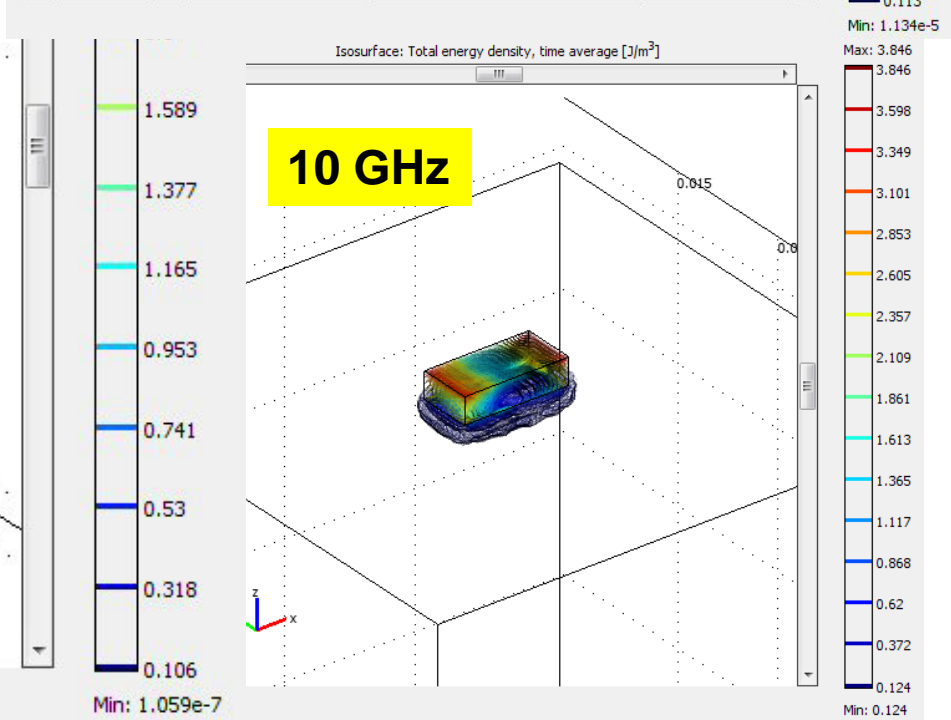
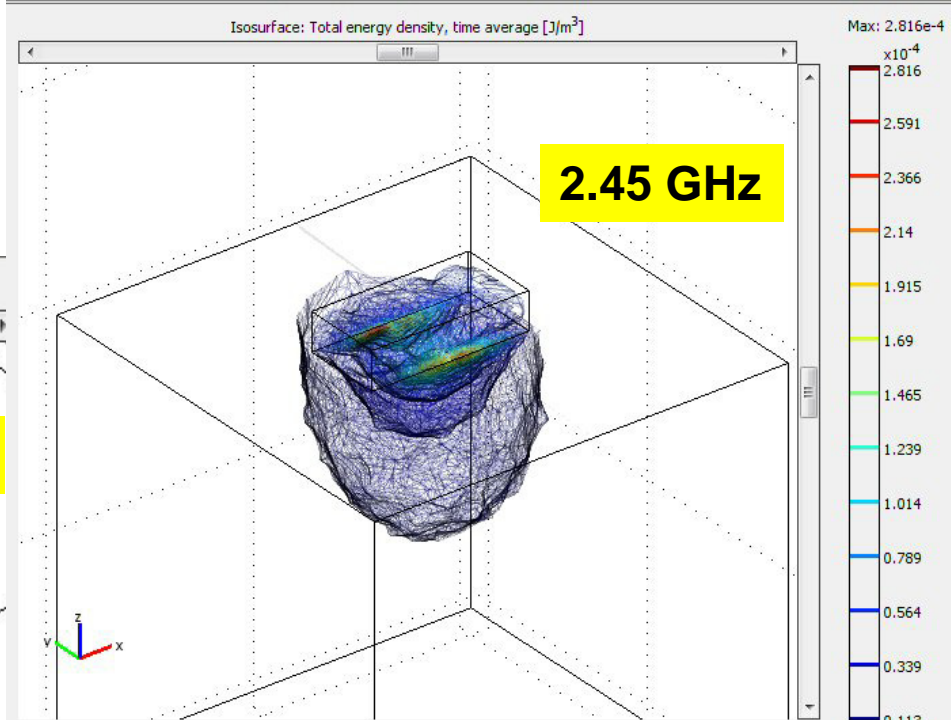
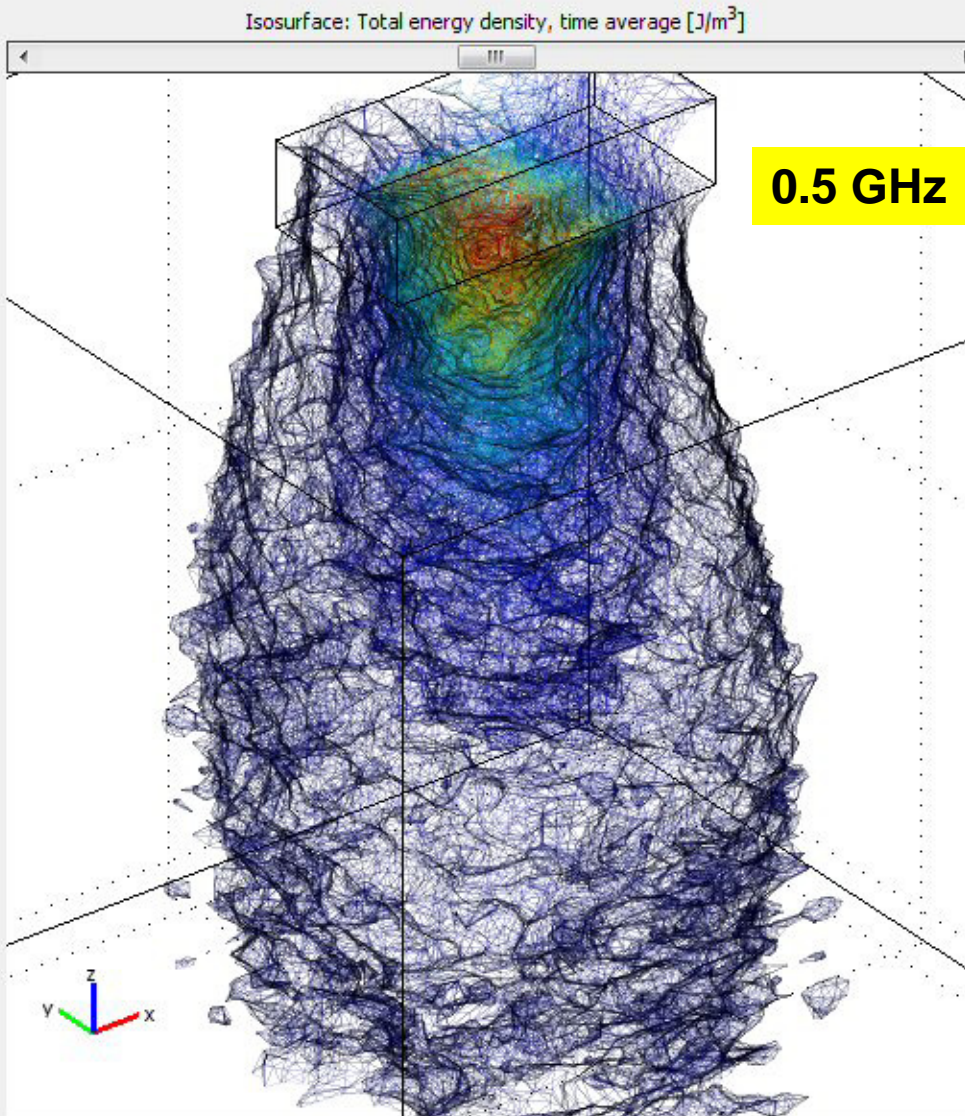


Three Microwave Frequencies

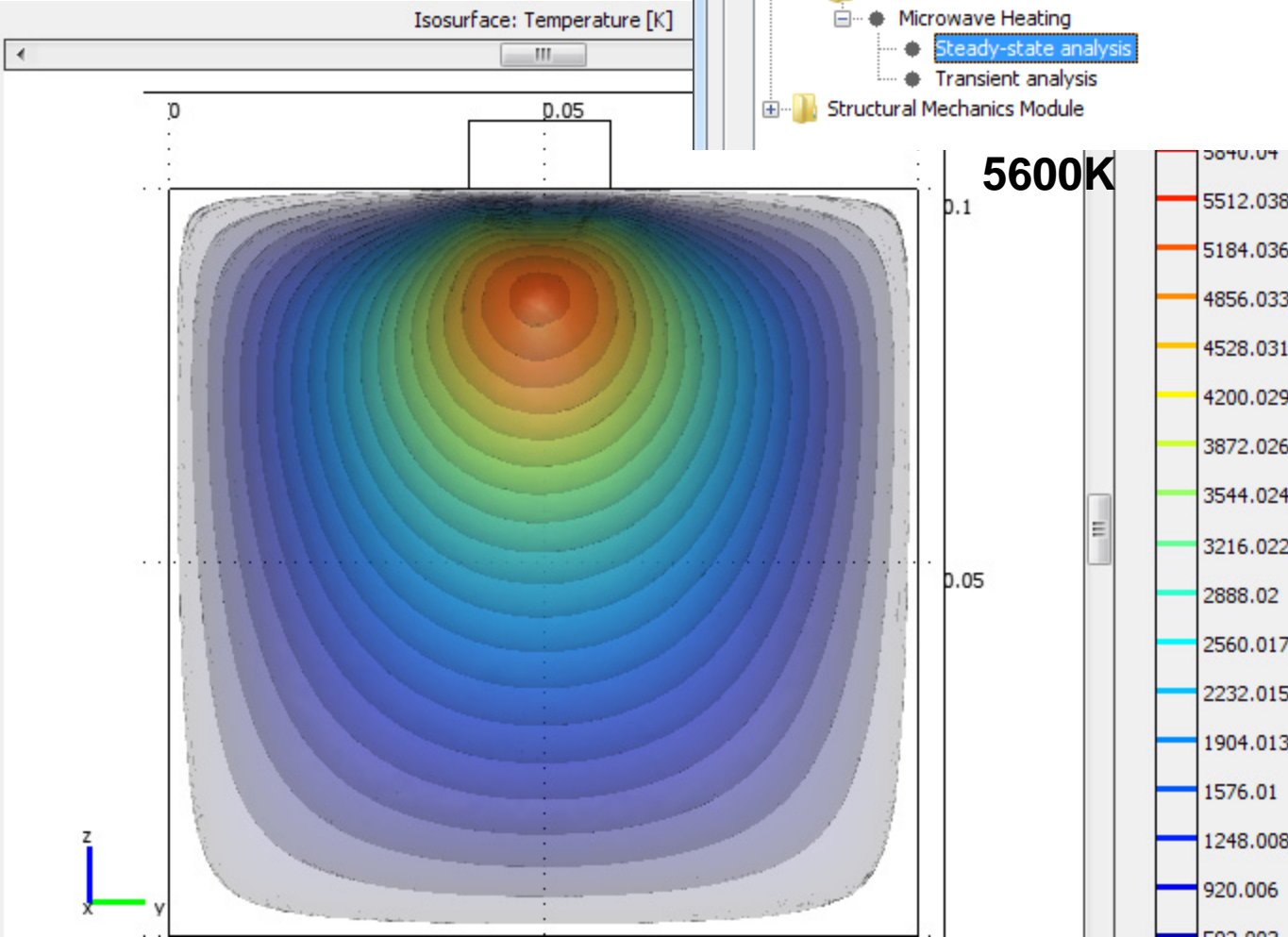
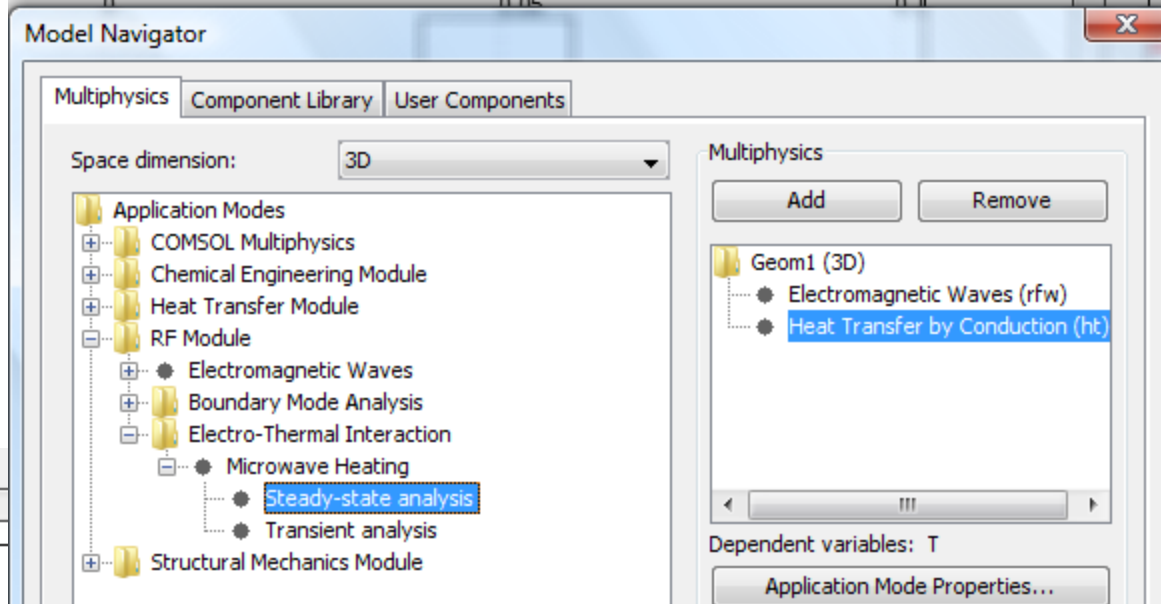


Microwave flanges for the three different microwave frequencies (0.9 GHz, 2.45 GHz and 10 GHz) used in this project showing the relative sizes of the experimental and test measurement hardware. Their standard sizes are designated WR975, WR340 and WR90 respectively. Each microwave frequency requires different geometry for COMSOL.

RFW -Total Energy Density Iso-Surface Penetration into Simulant

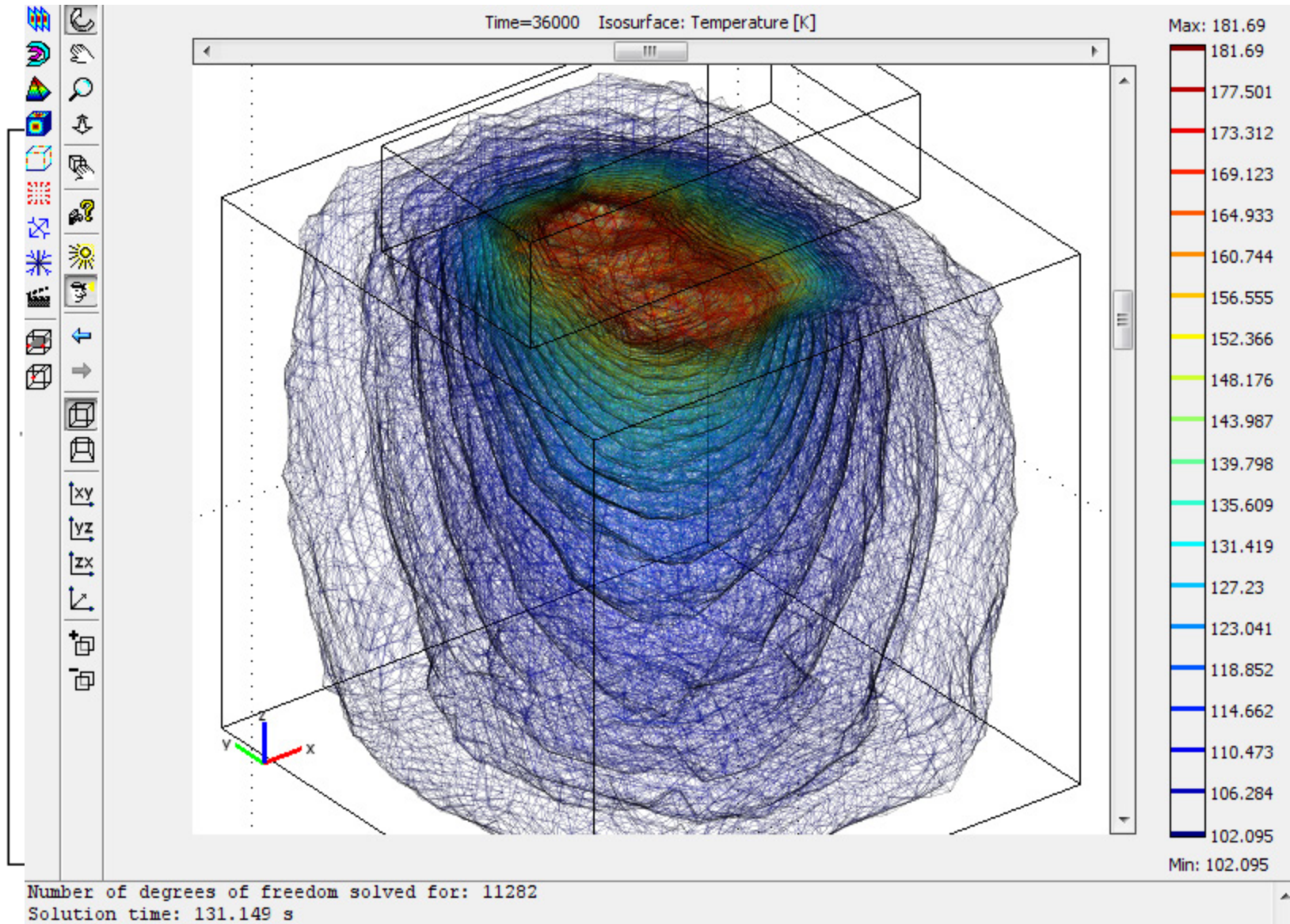


Steady-State Electromagnetism Coupled with Heat Transfer by Conduction

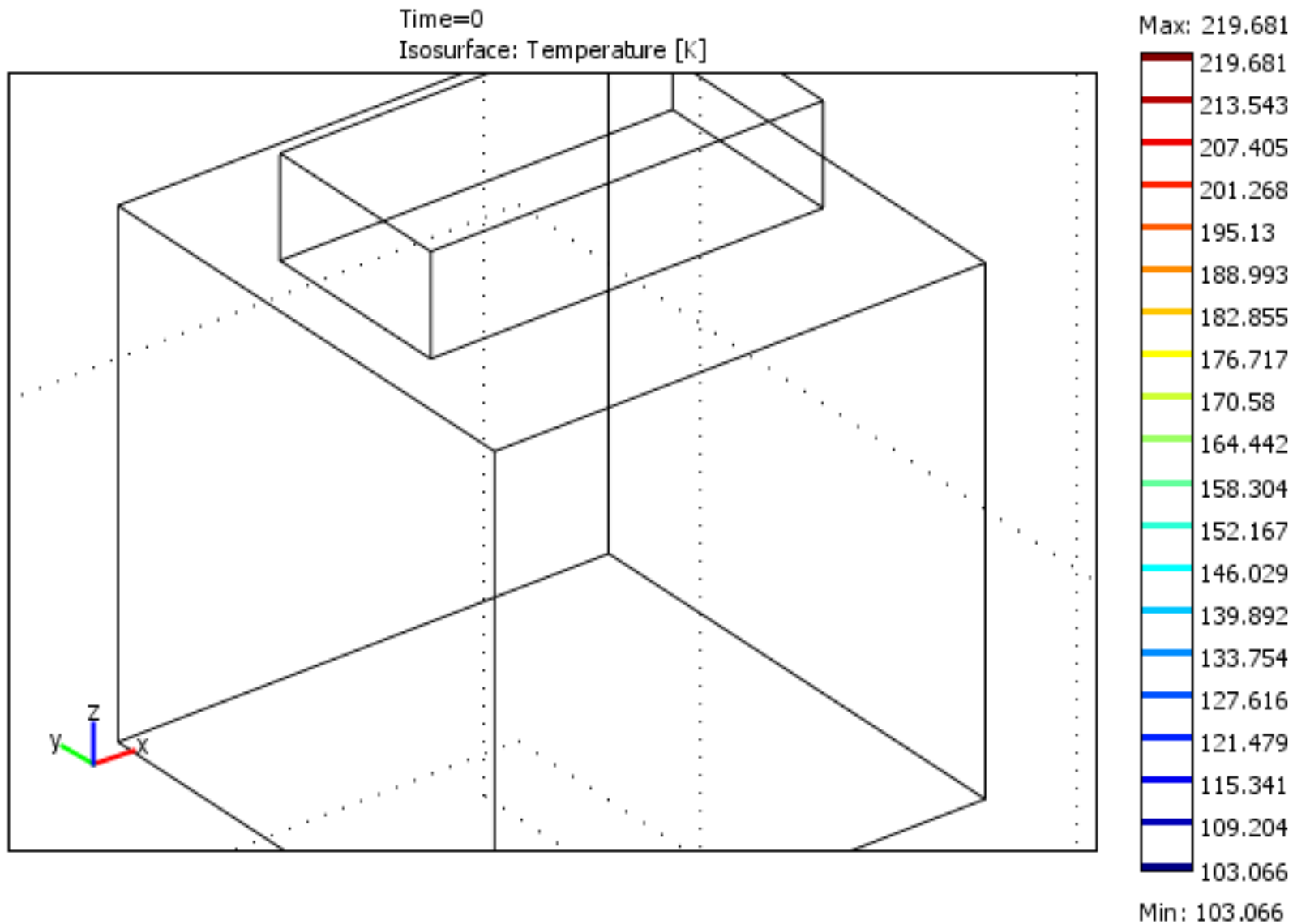


**2.45 GHz
1 KW power
Temperatures to
5000K**

RFW coupled with Heat Conduction - Transient Temperature Isotherms, 10 hours Click to start animation



RFW coupled with Heat Conduction - Transient Temperature Isotherms, 10 hours



Application of COMSOL

- **Processing parameters and hardware requirements for water extraction is a complex multi-physics problem.**
- **Microwave coupling to materials and heating is dependent on frequency and materials properties.**
- **Materials properties are a function of frequency and temperature.**
- **Can calculate microwave penetration and heating, with frequency and temperature dependent lunar soil dielectric properties.**
- **To Do – Model the Percolation of water vapor through the soil (porous media) characterized by the Darcy constant (currently being measured by Southern Research Institute).**
- **Parametric modeling will permit the evaluation of processing parameters most suitable for prototype hardware development, testing, and trade studies.**

Acknowledgements:

NASA HQ – Joint Science Mission Directorate (**SMD**) & Exploration Technology Development Program (**ETDP**) Lunar Advanced Science and Exploration Research (**LASER**) program

MSFC Management – Prior Seed Funding

Co-Investigators:

Dr. William Kaukler – University of Alabama-Huntsville
Hardware Instrumentation
Dielectric Property Test Configurations

Frank Hepburn – Materials & Processes, MSFC-NASA
Dielectric Property Measurements - Network Analyzer

COMSOL – Walter Frei -Technical assistance with the models

Contacts:

MSFC Public Affairs – Steve.Roy@msfc.nasa.gov

MSFC Technology Licensing – 256-544-5353

ed.ethridge@nasa.gov

END